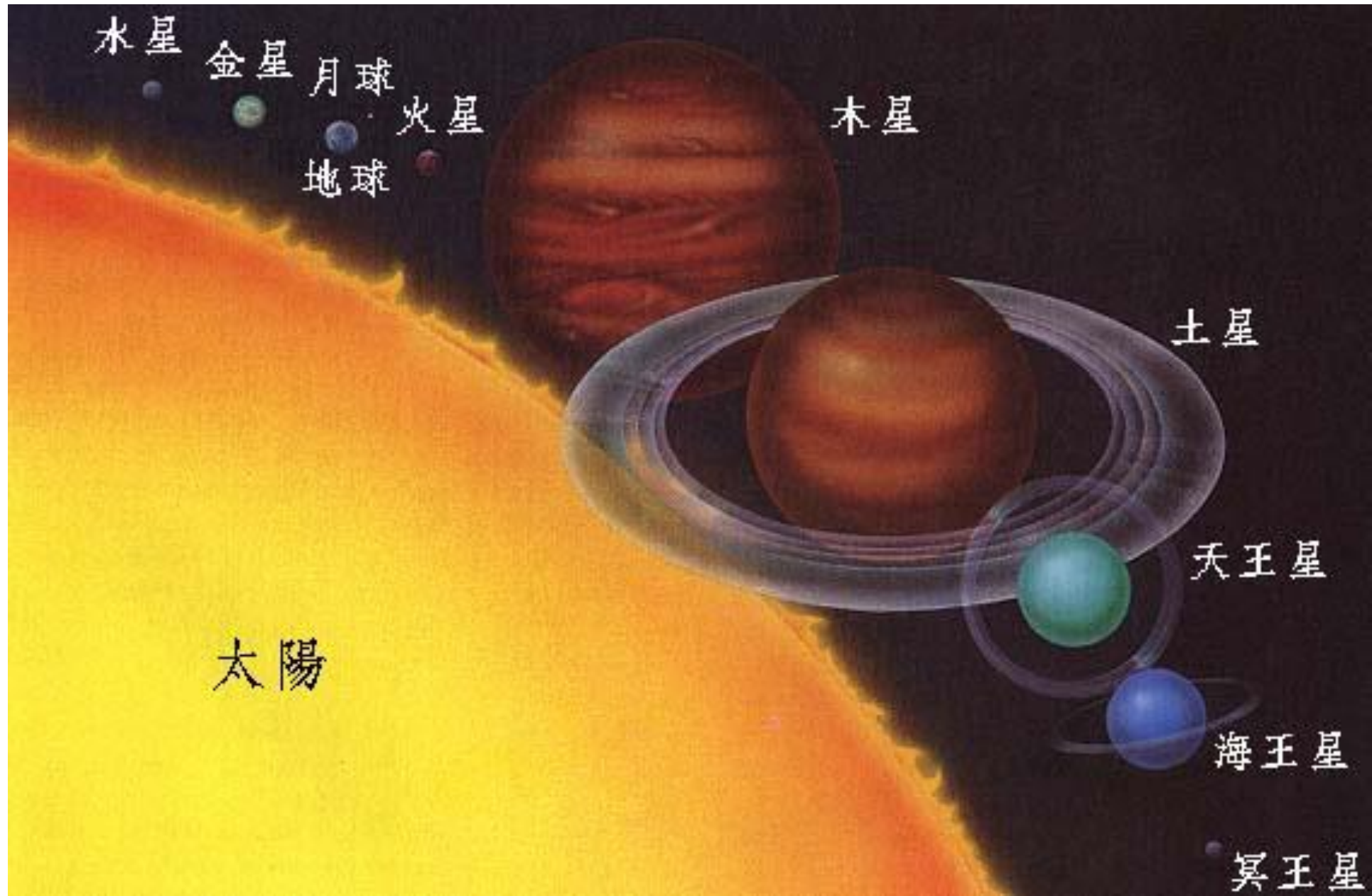
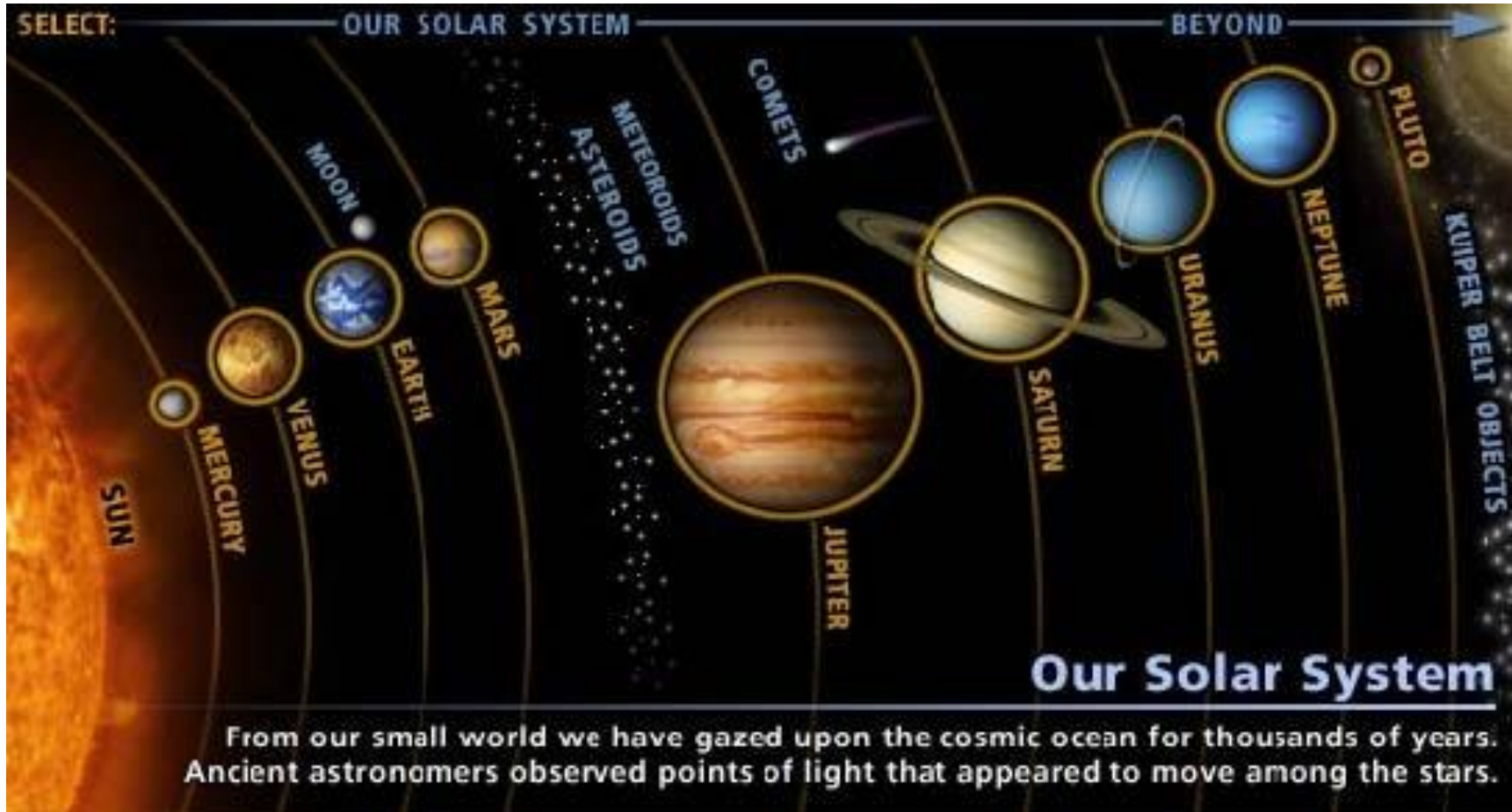


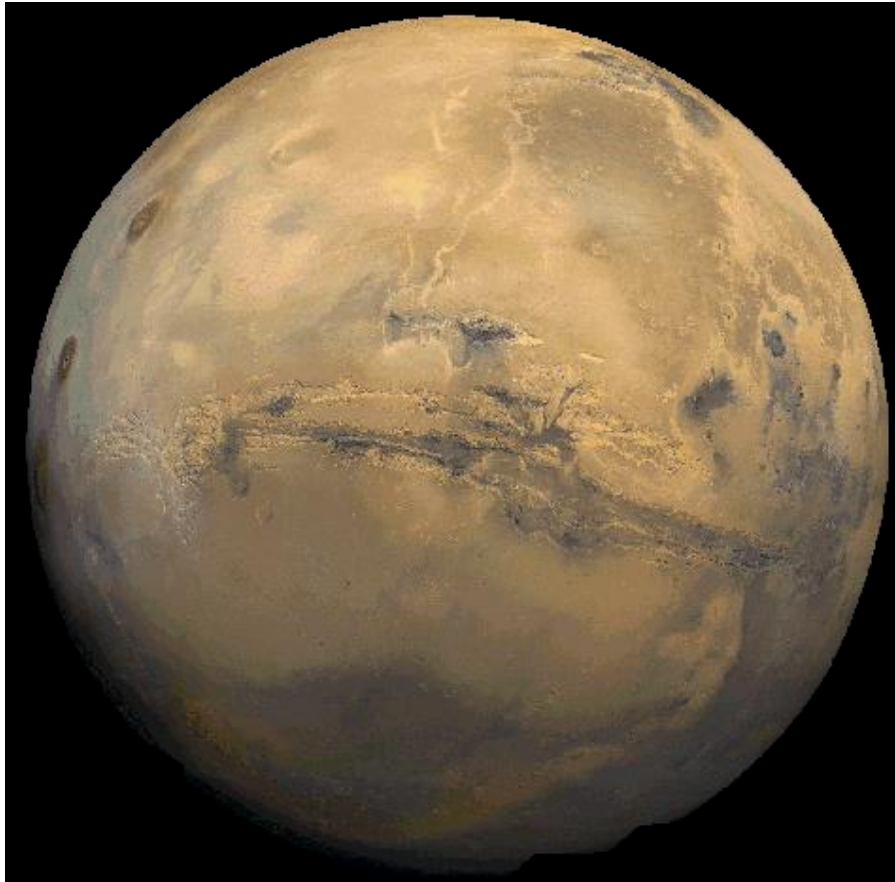
Interplanetary Space

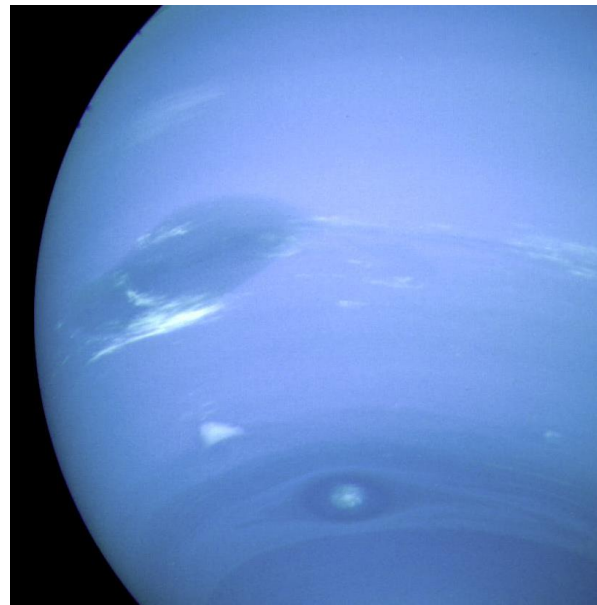
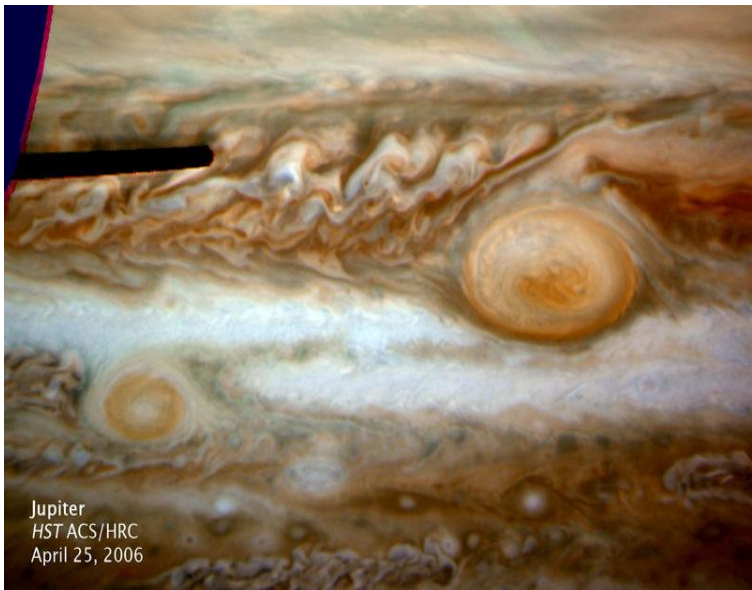
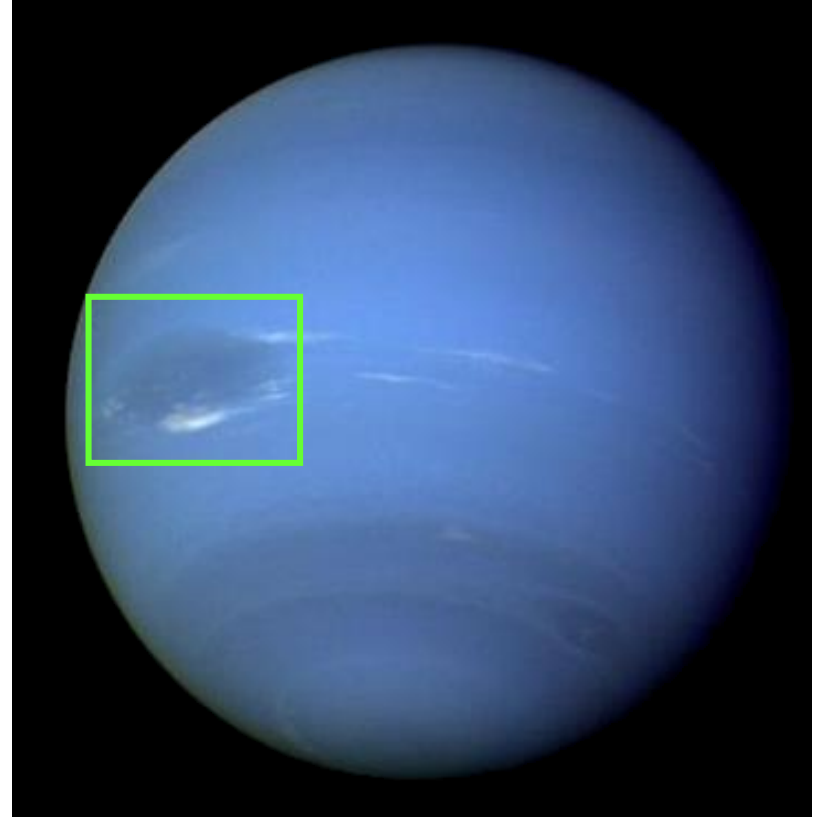
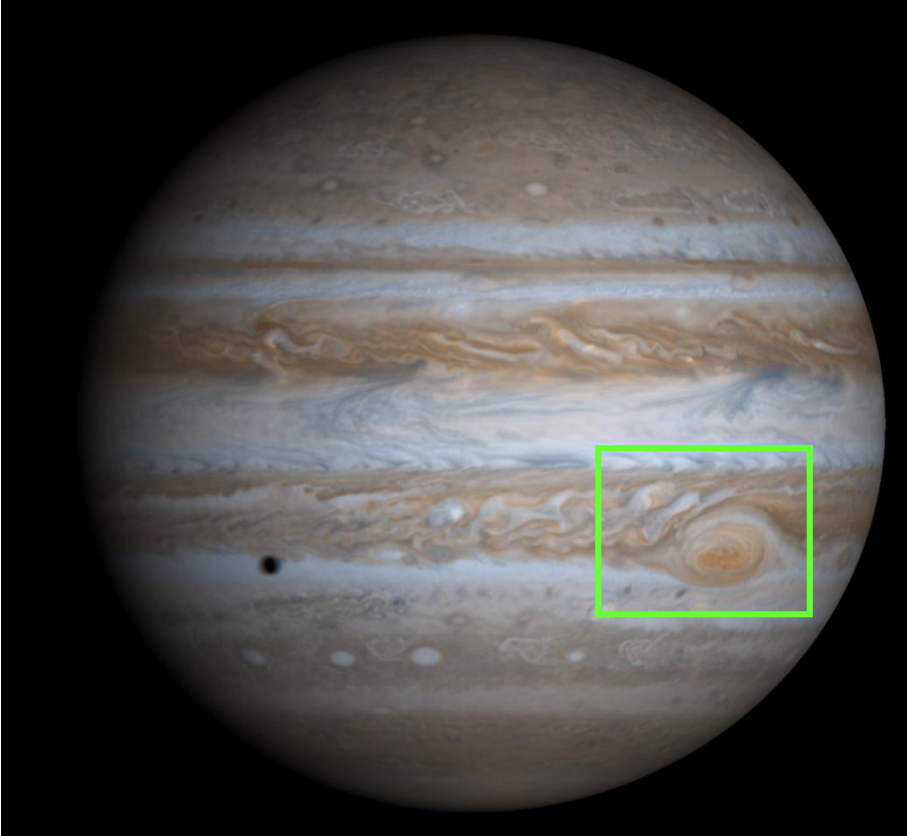
行星際太空

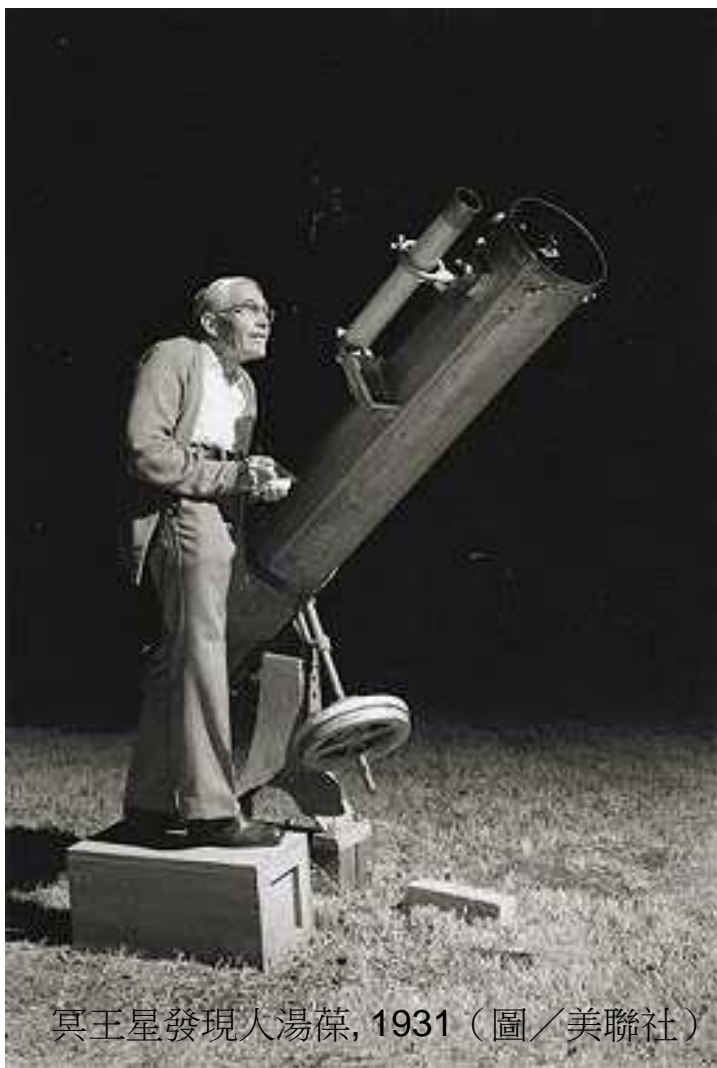




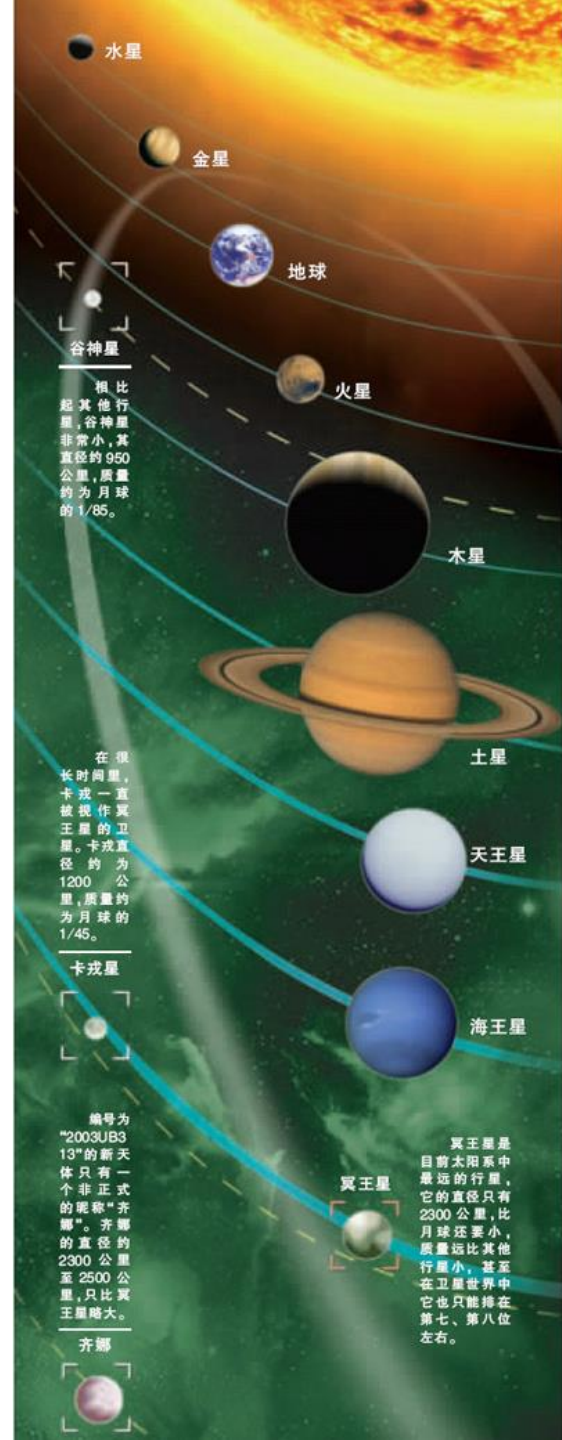
	表面溫度	自轉週期	公轉週期	衛星數	密度	重量	體積
水星	-170~430	58.65	88.97日	0	5.43	0.055	0.056
金星	420~485	243	245日	0	5.24	0.815	0.857
地球	-60~50	0.9973	365.26日	1	5.52	1	1
火星	-100~15	1.0260	686.98日	2	3.93	0.107	0.151
木星	-150	0.414	11.86年	61	1.33	317.83	1316
土星	-175	0.444	29.46年	31	0.69	95.16	745
天王星	-180	0.718	84.02年	25	1.27	14.54	65.2
海王星	-200	0.671	164.77年	13	1.64	17.15	56
冥王星	-215	6.38	247.8年	1	2.13	0.0022	0.006

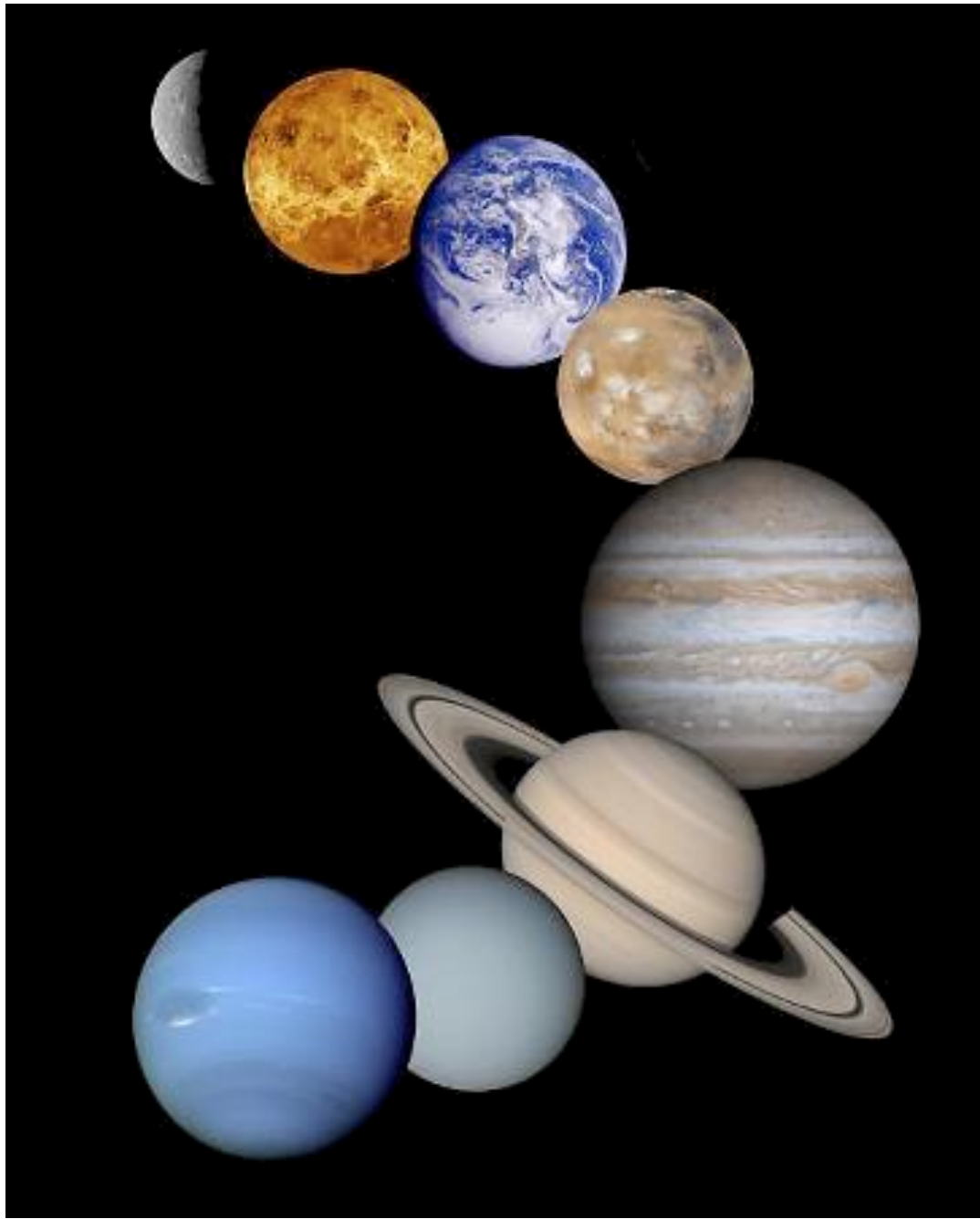




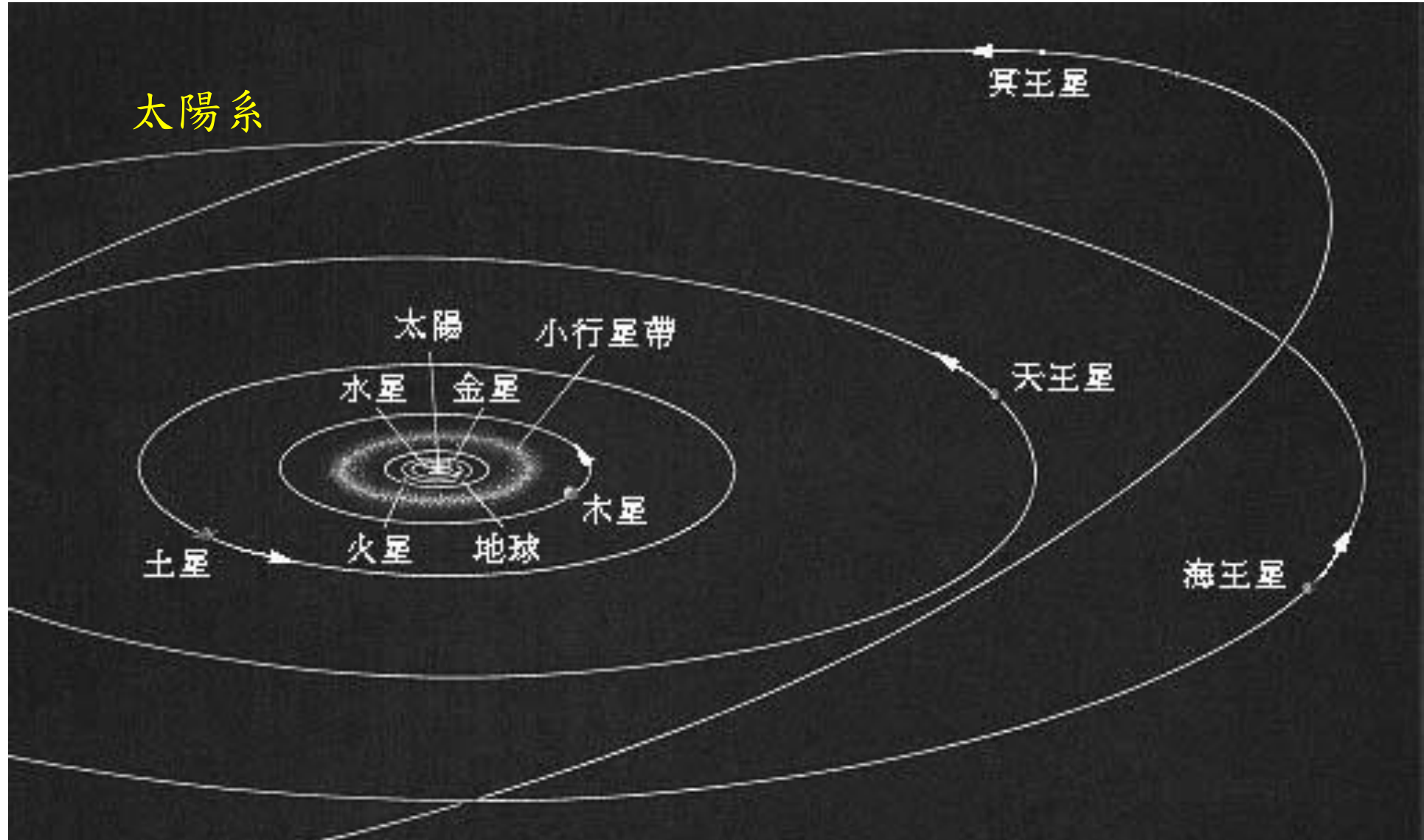


冥王星發現人湯葆, 1931 (圖/美聯社)





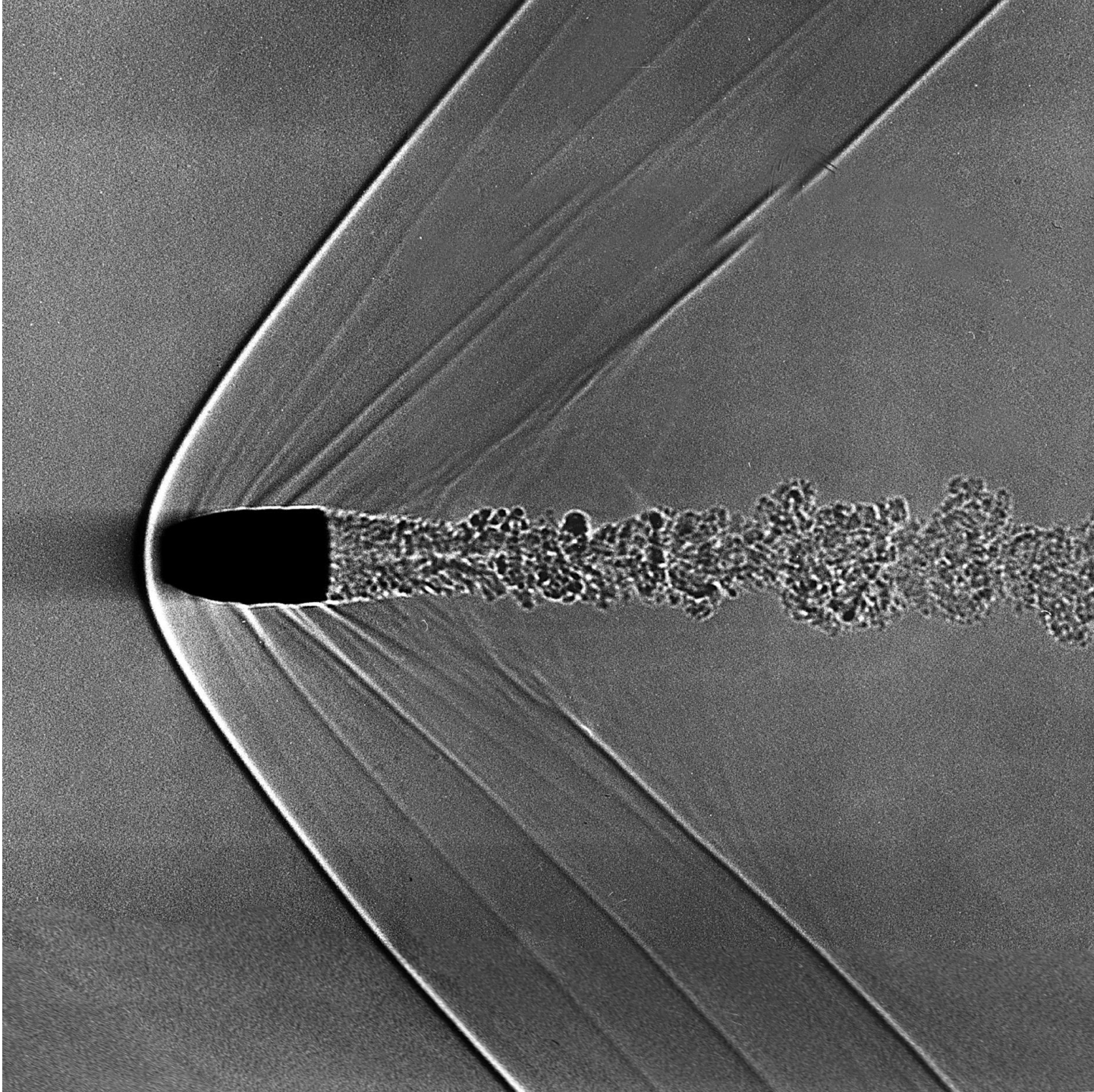
八大行星分別為水星、金星、地球、火星、木星、土星、天王星、海王星。

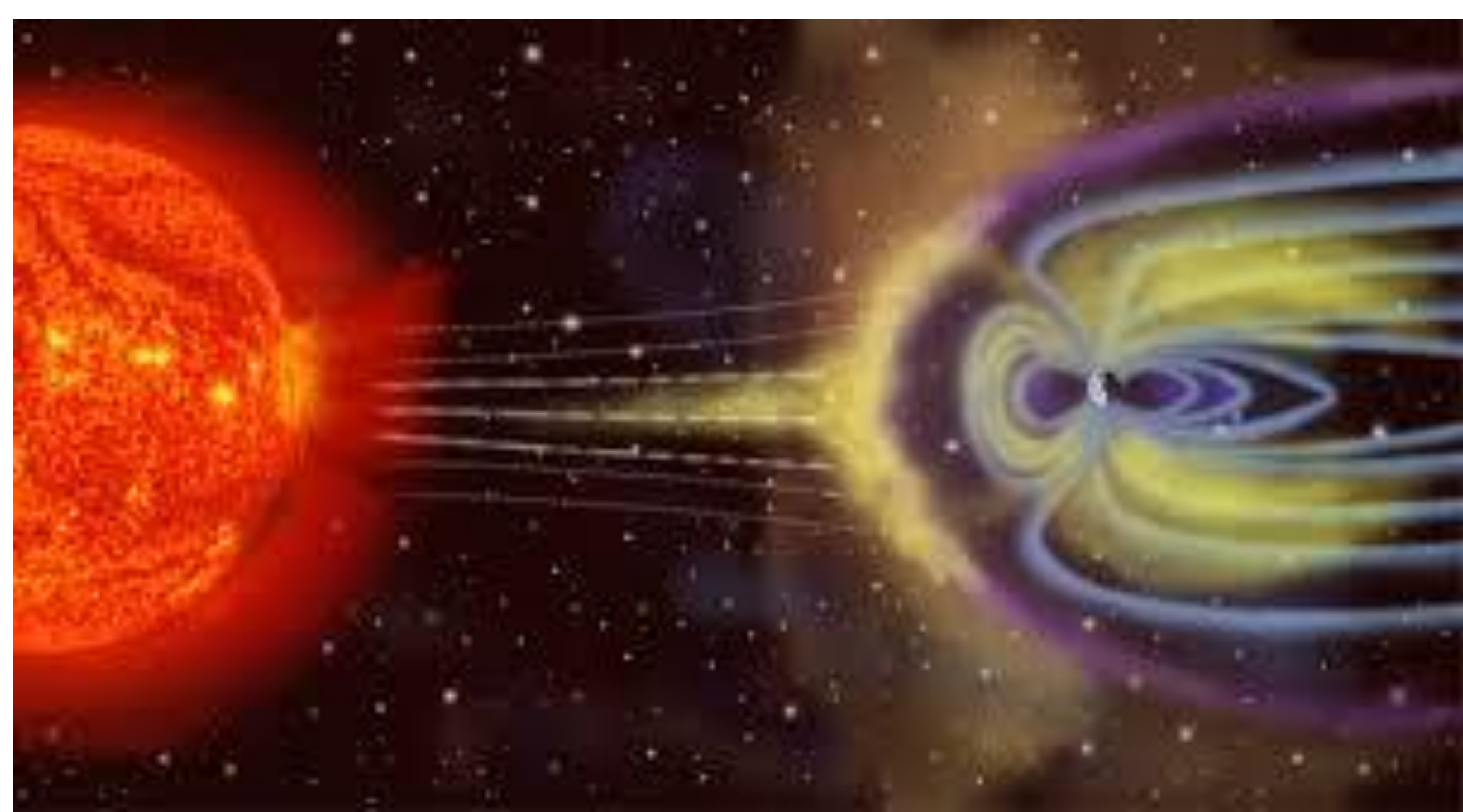


		大氣層	
		○	×
磁層	○	地球 木星 海王星	水星 天王星 土星
	×	金星 火星	

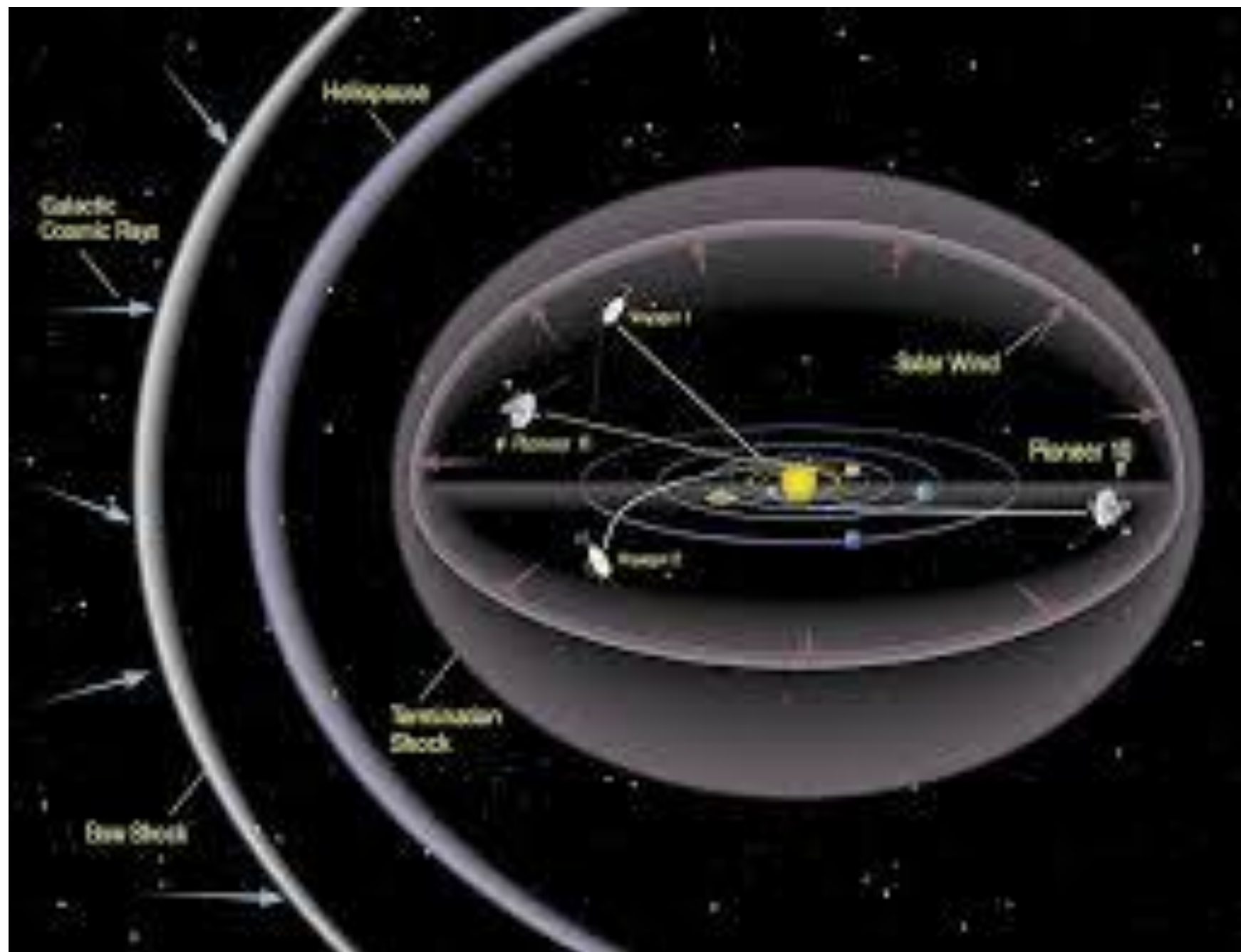
Shock





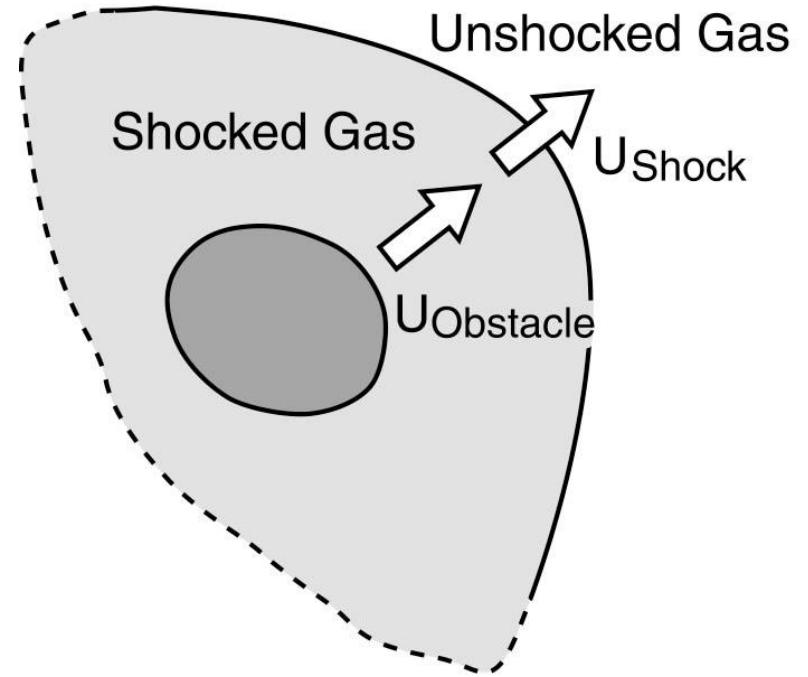




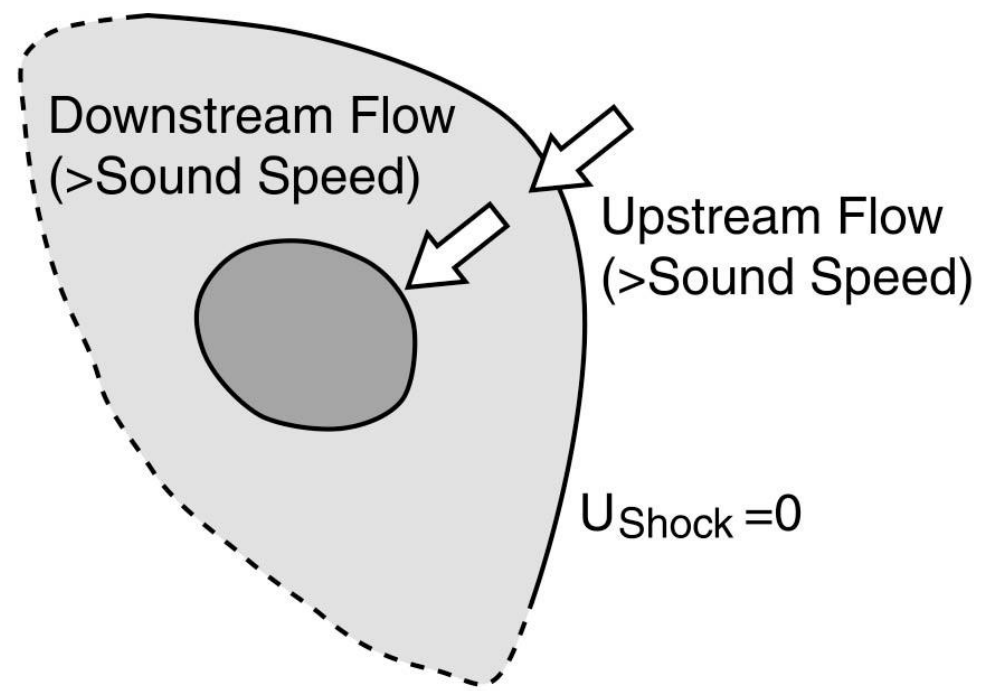


- A shock is a discontinuity separating two different regimes in a continuous media.
 - Shocks form when velocities exceed the signal speed in the medium.
 - A shock front separates the Mach cone of a supersonic jet from the undisturbed air.
- Characteristics of a shock :
 - The disturbance propagates faster than the signal speed. In gas the signal speed is the speed of sound, in space plasmas the signal speeds are the MHD wave speeds.
 - At the shock front the properties of the medium change abruptly. In a hydrodynamic shock, the pressure and density increase while in a MHD shock the plasma density and magnetic field strength increase.
 - Behind a shock front a transition back to the undisturbed medium must occur. Behind a gas-dynamic shock, density and pressure decrease, behind a MHD shock the plasma density and magnetic field strength decrease. If the decrease is fast a reverse shock occurs.
- A shock can be thought of as a non-linear wave propagating faster than the signal speed.
 - Information can be transferred by a propagating disturbance.
 - Shocks can be from a blast wave - waves generated in the corona.
 - Shocks can be driven by an object moving faster than the speed of sound.

- Shocks can form when an obstacle moves with respect to the unshocked gas.
- Shocks can form when a gas encounters an obstacle.



Moving Obstacle

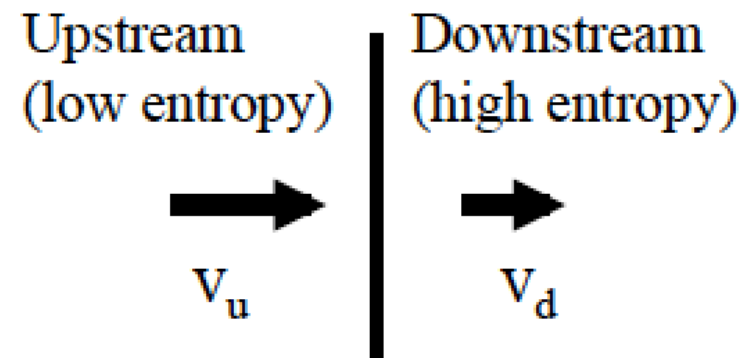


Stationary Shock

- The Shock's Rest Frame
 - In a frame moving with the shock the gas with the larger speed is on the left and gas with a smaller speed is on the right.
 - At the shock front irreversible processes lead to compression of the gas and a change in speed.
 - The low-entropy upstream side has high velocity.
 - The high-entropy downstream side has smaller velocity.
- Collisionless Shock Waves
 - In a gas-dynamic shock collisions provide the required dissipation.
 - In space plasmas the shocks are collision free.

- Microscopic Kinetic effects provide the dissipation.
- The magnetic field acts as a coupling device.
- MHD can be used to show how the bulk parameters change across the shock.

Shock Front



$$\begin{aligned}
 V_u &> V_d \\
 N_u &< N_d \\
 T_u &< T_d \\
 B_u &< B_d
 \end{aligned}$$